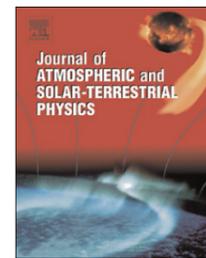




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Analysis of global Terrestrial Gamma Ray Flashes distribution and special focus on AGILE detections over South America



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ABSTRACT

Global distribution of the Terrestrial Gamma-ray Flashes (TGFs) detected by AGILE and RHESSI for the period from March 2009 to July 2012 has been analysed. A fourth TGF production region has been distinguished over the Pacific. It is confirmed that TGF occurrence follows the Intertropical Convergence Zone (ITCZ) seasonal migration and prefers afternoons. TGF/lightning ratio appears to be lower over America than other regions suggesting that meteorological regional differences are important for the TGF production. Diurnal cycle of TGFs peaks in the afternoon suggesting that Convective Available Potential Energy (CAPE) and convection are important for TGF production. Moreover all AGILE passages over South America in the same period have been analysed to find meteorological preferences for TGF occurrence. In each passage the analysis of Cloud Top Altitude (CTA), CAPE, number of strokes, number of storms and coverage area of clouds with temperatures below $-70\text{ }^{\circ}\text{C}$ (Cloud Top Coverage area, CTC) are computed. On average, AGILE has been exposed to 19,100 strokes between each TGF representing $\sim 68\text{ h}$ of exposure over active storms. High CAPE values, high cloud tops and high stroke occurrence suggest that meteorological conditions conducive to vigorous and electrically active storms are important for TGF production. It is shown that a high number of thunderstorms is preferable for TGF production which may be explained by a greater chance of the presence of a thunderstorm in the best development stage for TGF production. High tropopause altitude seems to be important but not primordial for TGF production.

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1. Introduction

CTR Wilson predicted in 1925 that electrons accelerated in high electric fields in thunderstorms would emit high energy radiation (Wilson, 1925). Many years later, this phenomenon was observed by Fishman et al. (1994), who reported 12 high energy emissions coming from Earth detected by the Burst and Transient Source Experiment (BATSE) on the Compton Gamma Ray Observatory (CGRO), with durations of few milliseconds and one or multiple peaks. This phenomenon, known nowadays as Terrestrial Gamma-Ray Flashes (TGFs), was first related to thunderstorms (Fishman et al., 1994). Then, TGFs were related to sferics, electromagnetic lightning emissions and sprites (Inan et al., 1996). Later, correlations of TGFs with positive polarity lightning (Cummer, 2005) and directly to positive intra-cloud (IC) flashes with peak currents between 15 kA and 57 kA (Stanley et al., 2006) refuted the theory

proposed by Inan (2005) that TGFs were produced by electromagnetic pulses radiated by high return strokes with high peak currents (450–700 kA). On the other hand, high energy photons detected from cloud-to-ground (CG) stepped leaders (Moore et al., 2001; Dwyer et al., 2003) and laboratory discharges (Dwyer, 2005; Dwyer et al., 2008; March and Montanyà, 2010, 2011) suggest that TGFs may be produced by similar processes. Analysis of the TGFs spectra had restricted the production altitude to 15–20 km (Dwyer and Smith, 2005; Carlson et al., 2007; Østgaard et al., 2008; Hazelton et al., 2009; Gjesteland et al., 2010). Moreover Lu et al. (2010) and Shao et al. (2010) find evidences of TGFs occurring during first stages of IC flashes. These results lead to conclude that Bremsstrahlung radiation emitted by runaway electrons during IC flashes is the most probable source of TGFs. The most accepted theories nowadays are the relativistic feedback mechanism (Dwyer et al., 2003; Dwyer, 2008, 2012) and lightning leader emission models (Dwyer, 2008, 2010; Carlson et al., 2009, 2010; Celestin and Pasko, 2011).

Other satellite instruments, mainly designed for the study of high energy astrophysics, have detected TGFs. These new

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instruments have allowed us to detect more events and better define their spectra and geographical occurrence. Smith et al. (2005) reported 86 TGFs detected by the Reuven Ramaty High Energy Spectroscopic Imager (RHESSI) (Smith et al., 2002) with energies up to 20 MeV. These TGFs showed that the geographical occurrence is greater over tropical continental regions in good agreement with the global lightning occurrence. Smith et al. (2010) estimated a global daily occurrence on Earth of 50 TGFs. Gjesteland et al. (2012) showed by an improvement of the search algorithm criteria for RHESSI measurements that the number of TGFs detected in the period 2004–2006 is twice the number previously reported. In 2006, Williams et al. (2006) ensured that IC flashes are the only candidates to produce TGFs because gamma rays produced at low altitudes by CG flashes would not be able to escape due to the atmospheric attenuation. This theory agrees with the higher TGF occurrence in the tropics where deep convection and high tropopause altitude allow high altitude thunderstorms and consequently occurrence of IC flashes higher in the atmosphere. On the other hand, Smith et al. (2010) showed using Monte Carlo simulations of gamma rays through atmosphere that tropopause altitude does not explain all the differences in the TGF geographical occurrence. Some other regional meteorological features may play an important role on these differences. Some parameters were studied by Splitt et al. (2010). He reported 805 TGFs observed by RHESSI with geographical occurrence in very good agreement with the three lightning continental tropical chimneys (Williams, 2005). Moreover, diurnal cycle and seasonal latitudinal distribution of TGFs and lightning are very similar in the tropics. Convective Available Potential Energy (CAPE) and outgoing infra-red flux comparison of the 805 TGF sub-satellite locations and 805 tropical random locations at the times of TGFs showed that TGFs occur in higher CAPE values than tropical background. Splitt et al. (2010) also reported the top altitudes and areal extent of 29 thunderstorms producing TGFs. The altitude varied between 13.6 and 17.3 km. The areal varied between 400 km² and 111,100 km² what showed that thunderstorms producing TGFs are not limited to larger convective systems. Finally a single case study of a thunderstorm producing a TGF was presented that showed an elevated mixed phase around 6 km.

The Mini-Calorimeter (MCAL) (Labanti et al., 2009) on-board of the Astrorivelatore Gamma ad Imagini Leggero (AGILE) satellite of the Italian Space Agency (ASI) (Tavani et al., 2009) was launched on 23 April 2007. First measurements reported a geographical distribution and diurnal cycle of TGFs consistent with RHESSI observations, and TGF photons with energies up to 40 MeV (Marisaldi et al., 2010). In 2011, Fuschino et al. (2011) demonstrated that TGF/lightning production rate was different in the three lightning chimneys, being greater over South America suggesting again that meteorological geographical differences play an important role in the TGF production. The most exceptional result of AGILE is its detection of photons up to 100 MeV (Tavani et al., 2011), whose spectrum is not consistent with Relativistic Runaway Electron Avalanche (RREA) (Gurevich et al., 1992; Dwyer and Smith, 2005). In 2014 the first official AGILE TGF catalog (Marisaldi et al., 2014) was published.

In this paper we present two different analysis. The first is a global occurrence comparison between RHESSI and AGILE TGFs of diurnal cycle and monthly distribution. The main objective of this analysis is to find regional differences in TGF production which may be explained by other meteorological factors than lightning altitude production. Previous publications suggested that lightning altitude production is important in TGF production but other meteorological factor may play an important role (Smith et al., 2010; Splitt et al., 2010; Fuschino et al., 2011). To enhance the study about meteorological geographical differences that may play an important role in TGF production, a second analysis has been

proposed. In this analysis different parameters that characterize atmospheric thunderstorm conditions are considered for all AGILE passages over South America during the period 2009–2012. The selected variables for the study are the Cloud Top Altitude (CTA), CAPE, number of strokes, number of storms and Coverage Area of clouds with temperatures below -70°C (Cloud Top Coverage, CTC). The objective is to identify most prevalent meteorological characteristics conducive to TGF production. The main purpose of this study is compare meteorological characteristics during passes with TGFs against those in passes in which TGFs were absent.

2. Data and methodology

2.1. Global analysis

2.1.1. TGF data

For the first analysis, geographical distribution, monthly rate distribution and diurnal cycle of RHESSI and AGILE TGFs in the period of March 2009–July 2012 have been compared, which is the period corresponding to the official AGILE catalog (Marisaldi et al., 2014). RHESSI is a National Aeronautics and Space Administration (NASA) spacecraft orbiting Earth in a low-Earth orbit (600 km) with an inclination of 38° launched 5 February 2002 designed to detect solar X and gamma rays. The instrument consists of nine Germanium detectors that can detect each photon in the range from 3 keV to 17 MeV individually without the need of an on-board trigger (Smith et al., 2002). In this paper we use the on-line catalogue of RHESSI TGFs provided by David M. Smith (<http://sprg.ssl.berkeley.edu/~dsmith>), with 283 TGFs in the period of March 2009–July 2012.

AGILE is a space mission of the Italian Space Agency with four different instruments to study gamma ray astrophysics. It was launched on 23 April 2007 in a low-Earth orbit at 550 km altitude with an inclination of 2.5° , therefore covering with high exposure a very narrow band across the equator. The main instrument used for TGF detection is the MCAL (Labanti et al., 2009) and is formed by 30 CsI(Tl) scintillator bars with an energy range from 300 keV to 100 MeV. We use the official AGILE TGF catalog (Marisaldi et al., 2014), 308 TGFs, with maximum photon energy 30 MeV, detected in the period of March 2009–July 2012, available at <http://www.asdc.asi.it/mcaltgfcatalog/>.

2.1.2. Lightning data

For the same period, lightning data of the Very Low Frequency lightning detection network called World Wide Lightning Location Network (WWLLN, Rodger et al., 2006) is used. In order to be compared with RHESSI and AGILE distributions WWLLN strokes have been analyzed in latitudinal ranges covered by both satellites, $\pm 44^{\circ}$ for RHESSI and $\pm 8^{\circ}$ for AGILE, the extra latitude interval with respect to satellite orbital inclination accounts for a maximum TGF satellite footprint distance in the order of 600 km. A relative WWLLN efficiency is provided by Hutchins et al. (2012) and used for the global longitudinal distribution. This detection efficiency is calculated from 14 April 2009 up to know. Therefore, the global longitudinal distribution plot presents RHESSI, AGILE and WWLLN from 14 April 2009 to 31 July 2012. For the latitudinal and diurnal cycle graphics data is plotted from March 2009 to July 2012. Finally, the monthly rate distribution for 3 years of data, from March 2009 to February 2012, has been analyzed.

2.2. AGILE passages

2.2.1. AGILE passages and GOES-East images

For the period from March 2009 to July 2012 a total of 15,951 passages over South America are considered. The data provided is

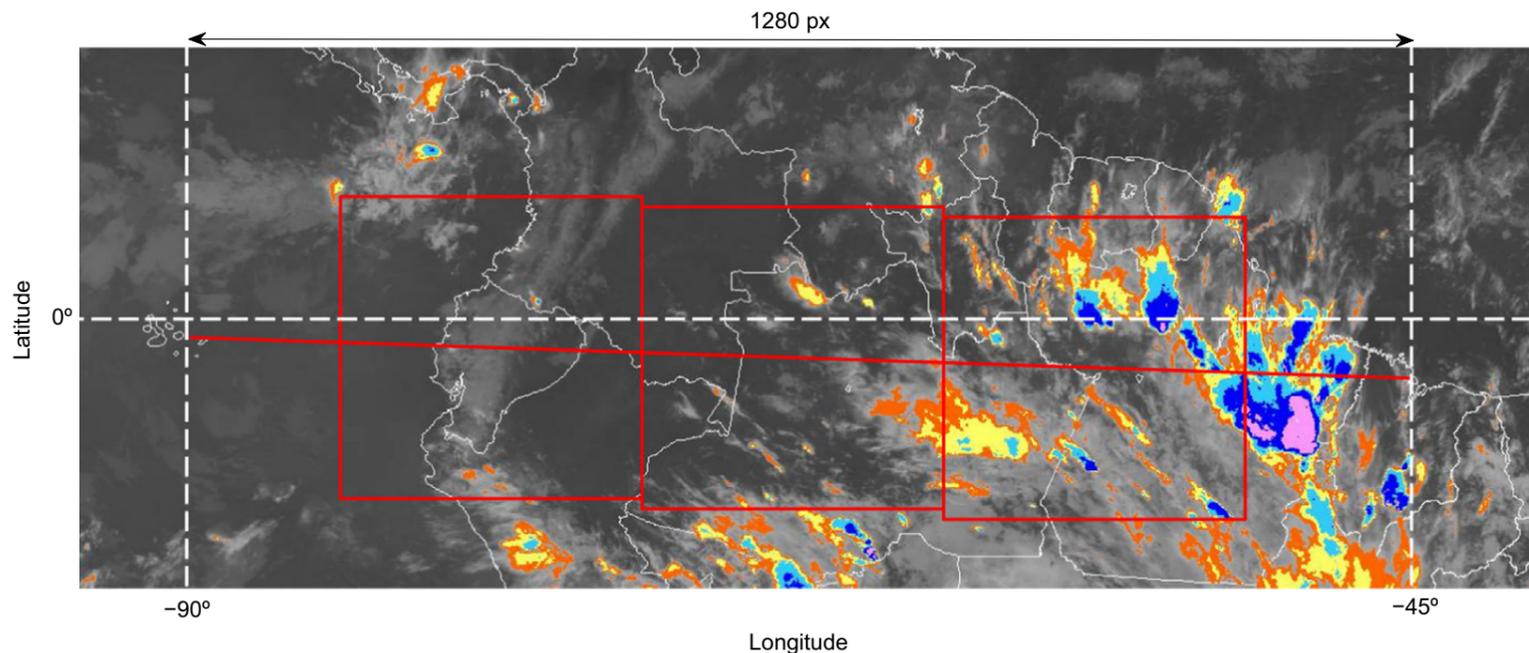


Fig. 1. Example of passage analysed. Red line is the satellite orbit and red squares are the squares of 280 pixels side (1230 km) used in the analysis. (For interpretation of the references to color in this figure caption, the reader is referred to the web version of this paper.)

a set of sub-satellite geographic coordinates and UTC time where t_1 corresponds to the satellite passage at -90° longitude and t_2 corresponds to -45° longitude for each passage. Infra-red images of channel 4 of the GOES-East satellite have been also used. GOES-East is a geosynchronous satellite at 0° latitude and -75° longitude at 35,800 km altitude that scans half the Earth every 15 min. For each AGILE pass over South America the corresponding GOES images were chosen. Taking four geographic reference points a function was used to transform passage geographic coordinates to GOES image pixels. Each AGILE passage is then ~ 1120 pixels wide in longitude. These 1120 pixels corresponds to -45° longitude ~ 4995 km (1° equals 111 km at the equator). That means that the GOES images used here have a pixel resolution of 4.4×4.4 km². The analysis of all the parameters is implemented in three squares of 1230 km side equidistant to -90° and -45° longitude. 1230 km side is the maximum distance of a point on the Earth surface from the AGILE subsatellite point that can be seen at the satellite altitude, hereafter called the field of view (FoV) for the purposes of the paper (Fig. 1). These squares are called AGILE FoV squares (Fig. 1). Knowing t_1 and t_2 and approximating AGILE track as a straight line, which is a reasonable assumption given the low orbital inclination, it is possible to calculate the time t_i ($i = 1, 2, 3$) where AGILE satellite is in the center of each square: $t_i = t_2 + (t_2 - t_1)/4i$, the total time t_t required by the satellite to cover the square, $t_t = (t_2 - t_1)/4$ and the times t_{eni} and t_{exi} that satellite enters and exits each square, $t_{eni} = t_i - t_t/2$ and $t_{exi} = t_i + t_t/2$. The analysis of each variable is then performed inside each AGILE FoV square. There are a total of 47,853 AGILE FoV squares without a TGF detected. For the 44 passages with a TGF detected it is taken only one AGILE FoV square. This one is centered at the subsatellite position when AGILE detected the TGF.

2.2.2. Cloud Top Altitude and coverage

Cloud analysis is done with GOES-East images of infra-red channel 4 ($10.20 \mu\text{m}$ – $11.20 \mu\text{m}$) provided by the Instituto Nacional de Pesquisas Espaciais (INPE) of Brazil (<http://satelite.cptec.inpe.br>). These images provide cloud top temperatures coloured in a range from -30°C to -80°C with 10°C intervals (Fig. 1). For each square the area covered by cloud tops with temperatures below -70°C is calculated by simply counting the pixels corresponding to these temperatures and multiplying by pixel area that is 4.4×4.4 km². If this area is larger than 0 it means that at least one high topped thunderstorm exists below the AGILE satellite. This calculation gives an approximation of the total area covered by

high-topped thunderstorms.

Because radiosonde stations are scarce in tropical regions, we consider models and reanalysis. MSISE-90 model is employed to infer Cloud Top Altitudes from temperatures given by GOES satellite infrared images (<http://ccmc.gsfc.nasa.gov/modelweb/atmos/msise.html>). For each square it is assumed that the atmosphere is in equilibrium. A vertical profile of neutral temperature is given by the MSISE-90 model at the center of the square at the time the satellite is passing over it, obtaining altitudes corresponding to GOES satellite infra-red image temperatures, that allows to identify maximum Cloud Top Altitudes in the square of thunderstorms when a TGF is detected or not.

There were only GOES-East images available for 8577 AGILE passages. Cloud Top Altitude and Coverage have been analysed for 25,731 AGILE FoV squares without TGF and 23 with TGF. For the analysis of the number of strokes, the number of thunderstorms and CAPE, GOES-East data is not needed.

2.2.3. Strokes and thunderstorms

Data from WWLLN is used to check lightning and thunderstorms. The total number of lightning strokes in the FoV of AGILE satellite when is over each square is the number of detections by WWLLN inside the square between times t_{eni} and t_{exi} . The mean value of $t_{exi} - t_{eni}$ is 3.2 min. To calculate the number of storms on each square, the total number of strokes inside each square in the period $t_{eni} - 20$ min and $t_{exi} + 20$ min has been found, clustering these strokes using a commercial software package (MATLAB R2012b, The MathWorks Inc., 2012) own function that calculates centroid distance of the cluster (Unweighted Pair Groups with Arithmetic Mean). The criteria values to build the storms are a maximum centroid distance of 1° and at least five strokes.

2.2.4. CAPE

CAPE data is obtained from European Centre for Medium-Range Weather Forecasts (ECMWF, <http://www.ecmwf.int/research/era/do/get/era-interim>), defined as “amount of potential energy an air parcels acquires when lifted adiabatically from its lifting condensation level to the level of neutral buoyancy”. This data is given every 3 h with $0.7^\circ \times 0.7^\circ$ resolution. We take the maximum CAPE before the satellite passage inside each square. This CAPE value will be between 3 h and 0 h before the satellite passage. Although it can be far from a TGF detection, previous CAPE better describes energetic conditions prior to storms producing TGF because CAPE is removed as storms restabilize the

environment.

3. Results and discussion

3.1. Global analysis

Geographic longitudinal distribution of 294 and 283 TGFs detected by AGILE and RHESSI respectively, and WWLLN strokes for the April 2009–July 2012 period are plotted in Fig. 2. The analysis takes into account the differences in detection efficiency of the network in different regions, depending on the sensors geographical density, as well as the variation of efficiency at different times, when new sensors are added to the network. The relative detection efficiency of WWLLN in the whole Earth is provided by WWLLN team (Hutchins et al., 2012). This relative detection efficiency for WWLLN is calculated on a 1 degree grid for each hour. This data has been used to normalize the stroke density in Fig. 2. It is easy to identify in Fig. 3 three TGF producing regions which are the tropical continental regions America, Africa and Asia. Furthermore, a fourth TGF region is identified over the Pacific. This region appeared previously in Collier et al. (2011) but it was not mentioned. Although this region is less active on TGF production than the other three, TGF and WWLLN distributions shows relevant maxima compared to other oceanic regions. TGF to lightning ratio is clearly lower over America compared to the other two continental regions for both AGILE and RHESSI. This result differs from Fuschino et al. (2011) (AGILE) and Smith et al. (2010) (RHESSI) who found a higher TGF/lightning ratio of TGFs over America compared to other regions. Fuschino et al. (2011) took only lightning data for AGILE inclination region, from -2.5° to 2.5° latitude, unlike here where lightning data is considered from -8° to 8° latitude, covering whole AGILE FoV. However, Fuschino et al. (2011) and Smith et al. (2010) compared data from the LIS (Lightning Imaging Sensor) and OTD (Optical Transient Detector) instruments that detect total lightning (CG and IC) while here we are using WWLLN data that mainly corresponds to highly energetic CG flashes.

It can also be observed in Fig. 2 that the four AGILE production regions have single narrow peaks unlike RHESSI, which has wider distributions with several peaks. This difference is due to the difference in satellite inclinations (Fig. 3). Because RHESSI scans a wider range of latitudes there is land at more longitudes while lightning and TGF have higher occurrence over land and coastal regions than far out over the sea. Fig. 3 also shows longitudinal

limits used in the range of latitudes for RHESSI analysis of latitudinal distribution, monthly distribution and diurnal cycle for the four regions separately. These limits correspond to the minima of distributions in Fig. 2. These limits are from -180° to -120° for the Pacific, from -120° to -35° for America, from -20° to 60° for Africa and from 80° to 180° for Asia.

Fig. 4 shows the monthly latitudinal distribution within the four TGF production regions in a box-plot graph for the 2009–2012 period (note that the box-plot graph does not take into account the total number TGFs, only median and dispersion over latitude). This box-plot graph is presented only for RHESSI TGFs, since AGILE inclination is too small for this kind of analysis. TGF detections follow the same latitudinal migration throughout the year as the ITCZ (Asnani, 1993), being in June–August in the north and in December–February in the south. This result was already reported by Splitt et al. (2010). Fig. 4 also shows the monthly latitudinal spread of lightning production, according to WWLLN data, in the four regions as well as the TGF production deficiency in North America reported by Smith et al. (2010). For the Pacific region, TGFs occur at a wide range of latitudes consistent with a larger spread in lightning detections. In America, when ITCZ is in the south, TGFs are detected in a narrower range of latitudes probably because RHESSI is affected by South Atlantic Anomaly (SAA). On the contrary, when the ITCZ is in the north TGFs are detected in a wider range of latitudes although there are few in North America. African TGF and lightning production is clearly limited north of the equator by the existence of the Sahara desert. In the SE Pacific region, a wide range of latitudes can be observed in all months, although for latitudes above 20° or below -20° there are less TGFs than expected in relation to lightning production. This could be associated with the lowering of tropopause height at higher latitudes (Smith et al., 2010). In the four TGF production regions, medians of the locations do not follow a symmetric distribution with respect to the equator, but are displaced to the north, the same as the ITCZ.

In Fig. 5 the diurnal cycle for AGILE TGFs, RHESSI TGFs and WWLLN strokes in each region is plotted for the March 2009–July 2012 period. In the four regions, AGILE, RHESSI and WWLLN stroke distributions agree with the Diurnal Tropical Cycle of Precipitation (DTCP) which is dominated by a maximum during the afternoon due to thunderstorms and a secondary maximum during night or early morning (Asnani, 1993). It is complicated to analyse the Pacific region due to the low number of TGFs. In the other three regions the RHESSI TGF distribution has two peaks, the largest peak in the afternoon and the smaller peak in the early morning.

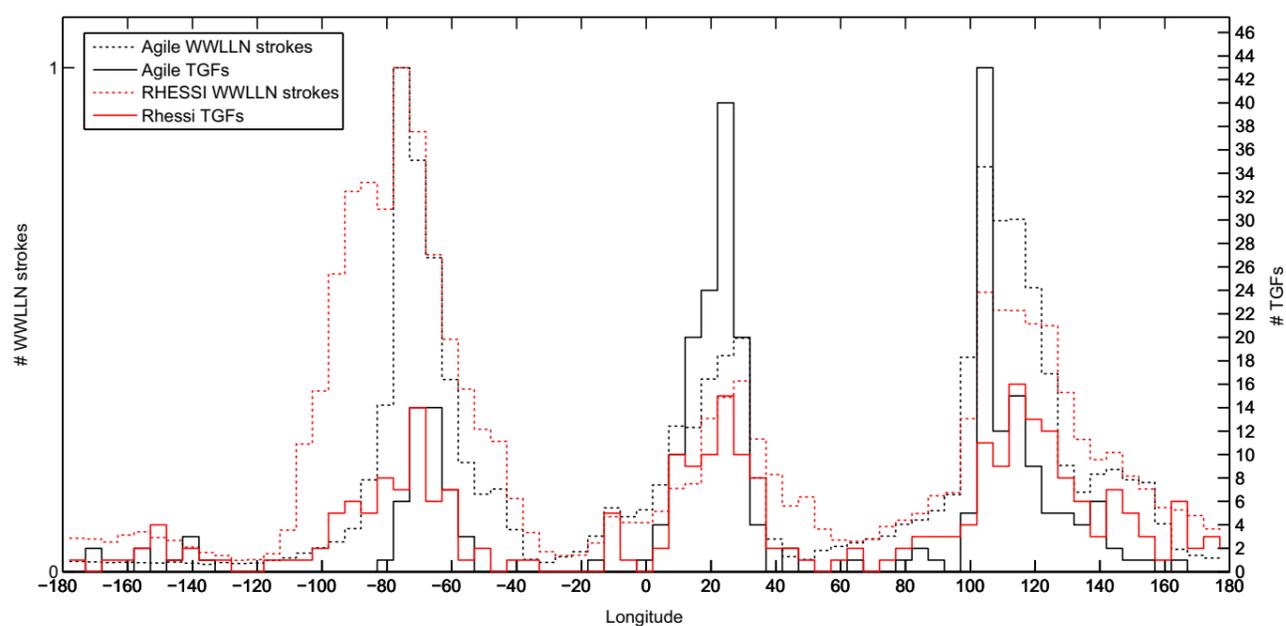


Fig. 2. Longitudinal distribution for April 2009–July 2012 period of AGILE WWLLN strokes (black dashed line), AGILE TGFs (black line), RHESSI WWLLN strokes (red dashed line) and RHESSI TGFs (red line). (For interpretation of the references to color in this figure caption, the reader is referred to the web version of this paper.)

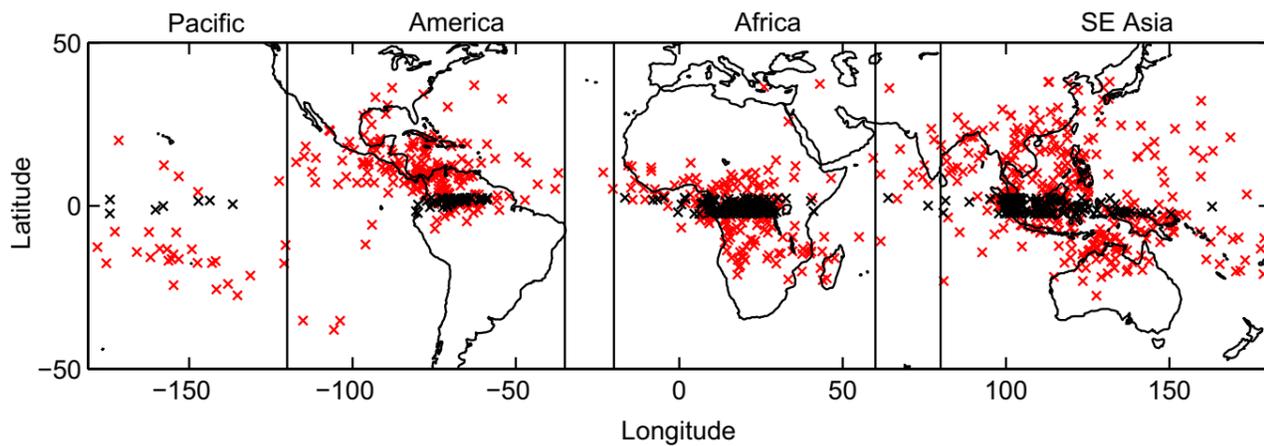


Fig. 3. Geographical distribution for March 2009–July 2012 period of RHESSI TGFs (red crosses) and AGILE TGFs (black crosses). Vertical black lines limits latitudes used for regional analysis. (For interpretation of the references to color in this figure caption, the reader is referred to the web version of this paper.)

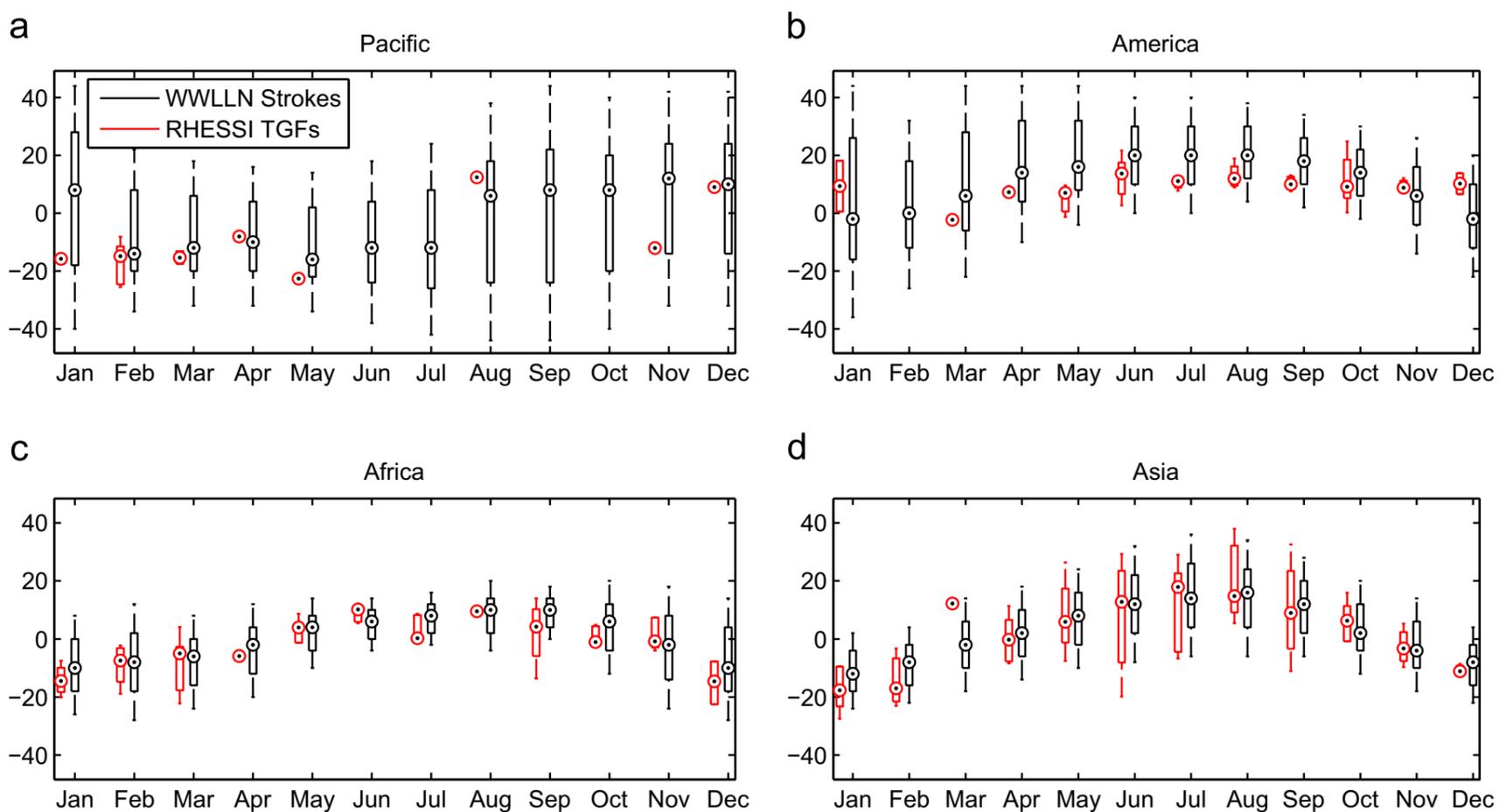


Fig. 4. Monthly boxplot graphic for March 2009–July 2012 period of latitudinal distribution for WWLLN strokes (black) and RHESSI TGFs (red). Circle is the median, box is the 50% of the data and whiskers around 95% of the data. (For interpretation of the references to color in this figure caption, the reader is referred to the web version of this paper.)

In America the peaks are very clear. The maximum in the afternoon coincides with the maximum of WWLLN strokes while the maximum in the night coincides with minimum of WWLLN strokes distribution. The same behaviour can be observed for RHESSI and WWLLN strokes distributions in Fig. 5 in Africa and Asia. Splitt et al. (2010) already reported the coincidence of this secondary peak of TGFs in the early morning with low lightning activity. The AGILE TGF distributions also show the predominant peak of TGFs during the night within the three subregions, which coincide with the predominant peak of WWLLN stroke distributions. In South America very few AGILE TGFs were observed during the night and early morning. Finally, in Africa there is an observable peak of AGILE TGFs in the night but is not coincident with the secondary maximum of WWLLN strokes at 2 LST. As will be seen in the following section, TGF production seems to be directly affected by the presence of large CAPE which in turn tends to be maximized the afternoons as a result of the solar heating.

The peaks March–May and September–November in Fig. 6 correspond to the ITCZ crossing the equatorial region covered by AGILE. RHESSI distributions are more flat due to the larger inclination of the satellite which in turn makes it less influenced by

the ITCZ movement. However over America during December–February TGF production almost drops to zero. During this season the ITCZ is south of the equator where TGF production should have been higher but satellite detection capabilities are affected by the South Atlantic Anomaly (SAA).

3.2. AGILE passages analysis

In this section the results for the meteorological conditions conducive to TGF production over South America are presented. The adopted method is the analysis and comparison of the meteorological conditions corresponding to AGILE passages related to TGF detections and those with not. Therefore there have been eliminated from the analysis all the FoV without TGF detections without storms. About 61% of total FoV without a TGF detected have at least one thunderstorm.

Forty-four TGFs have been observed for 83,8708 WWLLN strokes during a total exposure time of ~2550 h (this time is obtained by multiplying mean AGILE FoV time exposition, 3.2 min, by the total number of AGILE FoV squares). This corresponds 1 TGF detected by AGILE for 19,062 strokes observed by WWLLN during

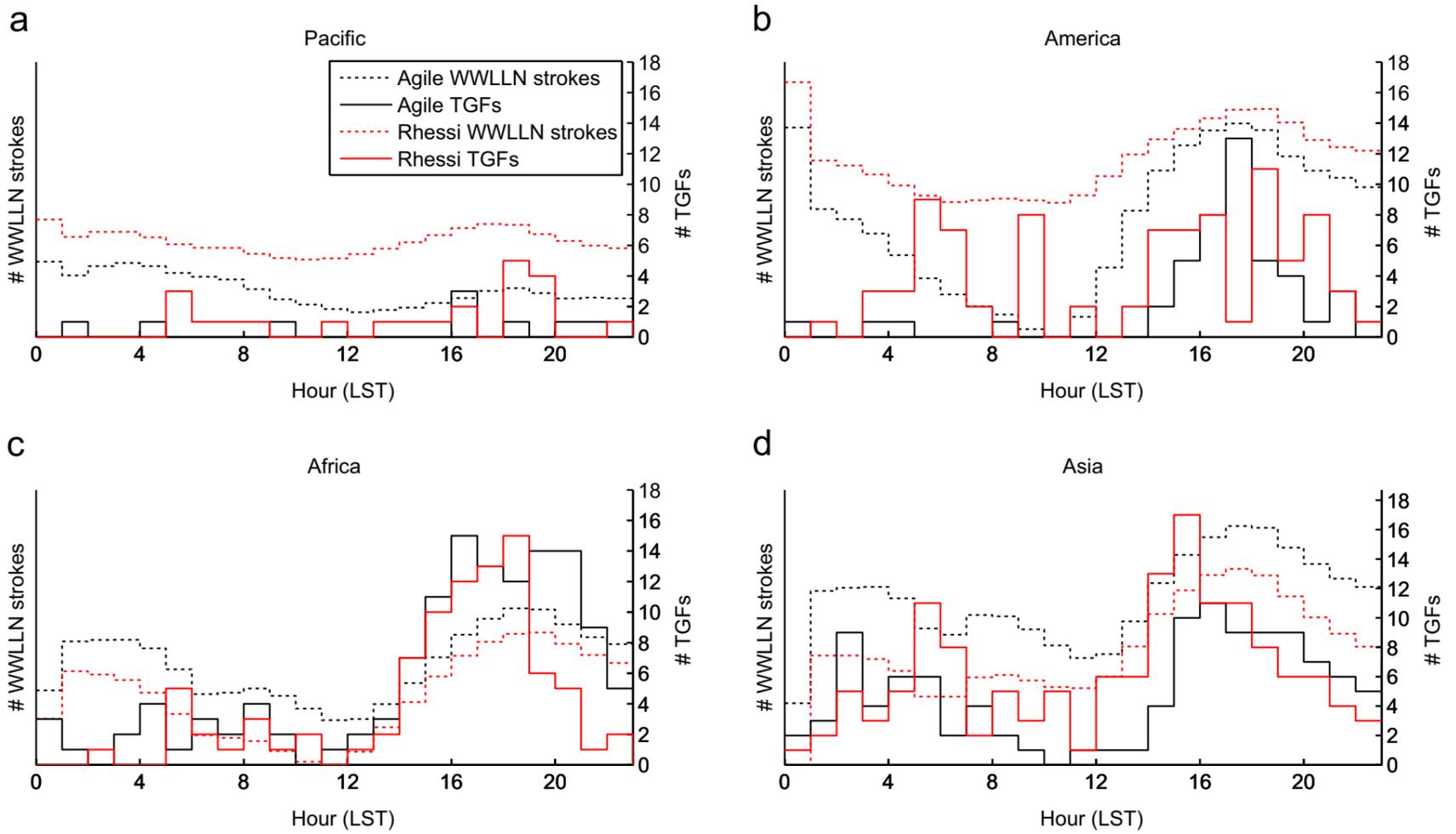


Fig. 5. Diurnal cycle for March 2009–July 2012 period for the four TGF production regions of AGILE WWLLN strokes (black dashed line), AGILE TGFs (black line), RHESSI WWLLN strokes (red dashed line) and RHESSI TGFs (red line). (For interpretation of the references to color in this figure caption, the reader is referred to the web version of this paper.)

~58 h of exposure. It is important to consider that global WWLLN efficiency is about ~10% (Virts et al., 2013) which means that AGILE is detecting ~1 TGF every 191,000 strokes and ~58 h of exposure. These ratios are remarkably lower than that reported in Fuschino

et al. (2011), based on the LIS/OTD data and correlation studies. We argue that the ratio reported here is underestimated because strokes are not weighted for the different detection probabilities given by the different distances to the satellite footprint. Stroke

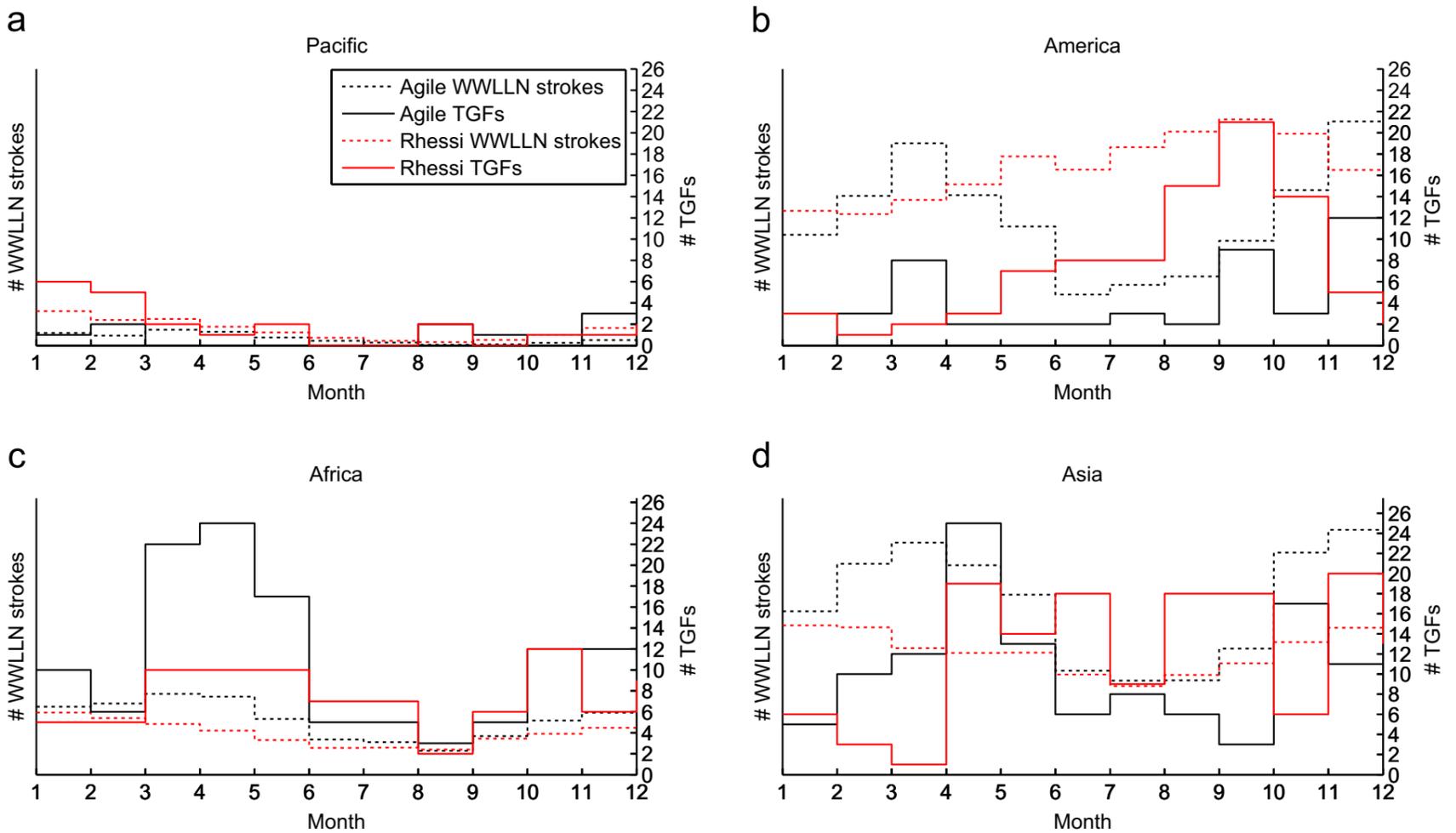


Fig. 6. Monthly distribution for March 2009–February 2012 period for the 4 TGF production regions of AGILE WWLLN strokes (black dashed line), AGILE TGFs (black line), RHESSI WWLLN strokes (red dashed line) and RHESSI TGFs (red line). (For interpretation of the references to color in this figure caption, the reader is referred to the web version of this paper.)

Table 1

Percentiles 10, 25, 50, 75 and 90 of the five meteorological parameters analyzed of AGILE passages detecting and non-detecting TGFs (non-detecting TGFs cases in parenthesis).

Percentile	TGF (non-TGF)				
	# Strokes	# Storms	CAPE (J/kg)	CTA (km)	CTC (km ²)
10%	12 (2)	5 (2)	0(0)	14(13)	21,304 (232)
25%	27 (5)	9 (3)	2588 (1606)	15 (13)	26,746 (639)
50%	59 (13)	14 (5)	3444 (2371)	16 (16)	46,774 (2749)
75%	205 (34)	20 (8)	4008 (3197)	17 (17)	57,422 (10,667)
90%	277 (70)	26 (13)	4651 (4064)	17 (17)	91,972 (27,654)

occurrence (flash km² yr⁻¹) is two orders of magnitude higher in equatorial South America than in Europe (Abarca et al., 2010). That means that same satellite over Europe would need ~5800 h (~242 days) of exposure to observe a TGF, neglecting the further contribution of meteorological factors. This analysis suggests that a very large effective area is a crucial requirement for space instruments designed to detect TGFs above Europe.

Table 1 shows the percentiles 10, 25, 50, 75 and 90 for the five meteorological parameters during AGILE passes with and without detected TGFs analysed: number of strokes, number of storms, Cloud Top Area (CTA), Convective Available Potential Energy (CAPE) and Cloud Top Coverage (CTC). Table 2 shows percentiles of the same five meteorological parameters for the AGILE passages without detected TGFs corresponding to percentile values of AGILE passages detecting TGFs of Table 1. For example, according to Table 1, the 25% percentile for the number of strokes in the case of TGF corresponds to 27. This number corresponds to the 69% percentile in the case of passages without TGF detection, see Table 2. Figs. 7–11 show diurnal cycle and monthly distributions of the same meteorological parameters for passages non-detecting TGFs. These distributions are plotted in stacked bar graphs of four categories. These categories are four ranges corresponding to quartiles and median values of detecting TGF passages. In this way we can figure out if diurnal cycle and monthly distribution of non-detecting TGF passages agrees with detecting TGF distributions of Figs. 5 and 6. Since TGFs have a clear diurnal and monthly signature (Figs. 5 and 6), it is important to assess this behavior also for non-detecting TGF passages.

3.2.1. Number of strokes and storms

About 75% of TGFs were detected when the AGILE satellite was exposed to more than 80 strokes/min (27 strokes in 3.2 min, time over AGILE FoV square, and WWLLN efficiency ~10% Virts et al., 2013) and nine active storms (Table 1). This is because of the more strokes and thunderstorms the higher is the probability to detect a TGF. Table 2 shows that these situations corresponding to 75% TGFs detected represents only less than 31% of AGILE passages non-detecting TGFs, meaning that TGFs are detected when satellite is exposed to higher strokes and storms rates. Moreover, in

Table 2

Percentiles of the five meteorological parameters analyzed of the AGILE passages non-detecting TGFs corresponding to the percentile values 10, 25, 50, 75 and 90 of AGILE passages detecting TGFs.

TGF percentile (%)	Non-TGF percentiles (%)				
	# Strokes	# Storms	CAPE	CTA	CTC
10	49	46	0	31	87
25	69	76	57	38	90
50	87	91	81	40	97
75	99	98	89	64	97
90	99	99	95	64	99

Figs. 7 and 8 it is shown that the number of strokes and the number of storms of AGILE passages non-detecting TGFs corresponding to 75% values of TGFs detected, peaks clearly in the afternoon as TGFs detected do. For monthly distribution also exists coincidence for the same values as diurnal cycle, although peaks are less pronounced.

3.2.2. CAPE

Table 1 shows that 75% of TGFs detected occurred when CAPE was higher than 2588 J/kg. This range corresponds to 43% of AGILE passages without detected TGFs. Fig. 9 shows that passages with CAPE values higher than 2588 J/kg tend to occur mostly in the afternoon without a monthly preference. These results reinforce what it was shown in the previous section, i.e. that TGFs mostly occurs at afternoon when solar irradiance results in high CAPE. So solar irradiance and CAPE plays an important role in helping to get the right conditions in thunderstorms for the TGF production and could explain the differences of regional TGF/lightning ratio presented previously. It is also worth noting that the largest CAPE values are observed at 15 LST, with a distribution very similar to that corresponding to passages with TGF detection. However, the peak in TGF detection is 2 h later, at 17 LST. This delay is possibly due to thunderstorm lightning activity evolution driven by high CAPE values. In fact, as shown in Fig. 5, lightning activity over South America peaks at 17 LST.

3.2.3. CTA

About 50% of TGFs occurred when CTA was >16 km (Table 1). On the other hand, CTA >16 km represents a 60 % of total AGILE passages (Table 1) which means that this altitude is quite typical in the tropics. In the analysis of diurnal cycle and monthly distribution in Fig. 10 it can be seen that there is a peak at noon which correspond to maximum solar irradiance. Williams et al. (2006) pointed out that TGFs produced high in the atmosphere experience less atmospheric attenuation and can be more easily detected at satellite altitudes. The results obtained in this analysis show clearly that high altitude thunderstorms are quite typical in the tropics where TGFs are more detected. If CTA was critical on TGF production more TGF would be expected due that most of thunderstorms reach altitudes ~16 km. We cannot confirm that the altitude is primordial for TGF detection. However it seems an important factor (note that almost all TGFs are detected when CTA over 14 km) but together with high CAPE and lightning activity.

3.2.4. CTC

For the CTC it can be seen in Table 1 that 75% of the detected TGFs occur when high cloud tops with temperatures below -70  C cover more than 26,000 km². These cases represent only 10% (Table 2) of the total AGILE passages non-detecting TGFs analyzed, so TGFs are detected under not typical conditions. Moreover, these cases occurs more often in the afternoon and during March–May and September–December periods (Fig. 11). These distributions are coincident with TGF diurnal cycle and monthly distribution (Figs. 5 and 6).

3.2.5. Summary

It is shown that 75% TGFs detected by AGILE over South America in the period from March 2009 to July 2012 prefer CAPE >2588 J/kg, stroke rates >80 strokes/min, number of storms >9 and CTC >26,000 km² (Table 1). These values are typical of tropical thunderstorms slightly displaced from ITCZ (Cooray, 2003, Chapter 1) which is consistent with seasonal migration of AGILE TGFs detected over SA (Figs. 5 and 6) and AGILE passes non-detecting TGFs (Figs. 7–11) as well. These high CAPE values >2000 J/kg are related to vigorous updrafts which are important for charge

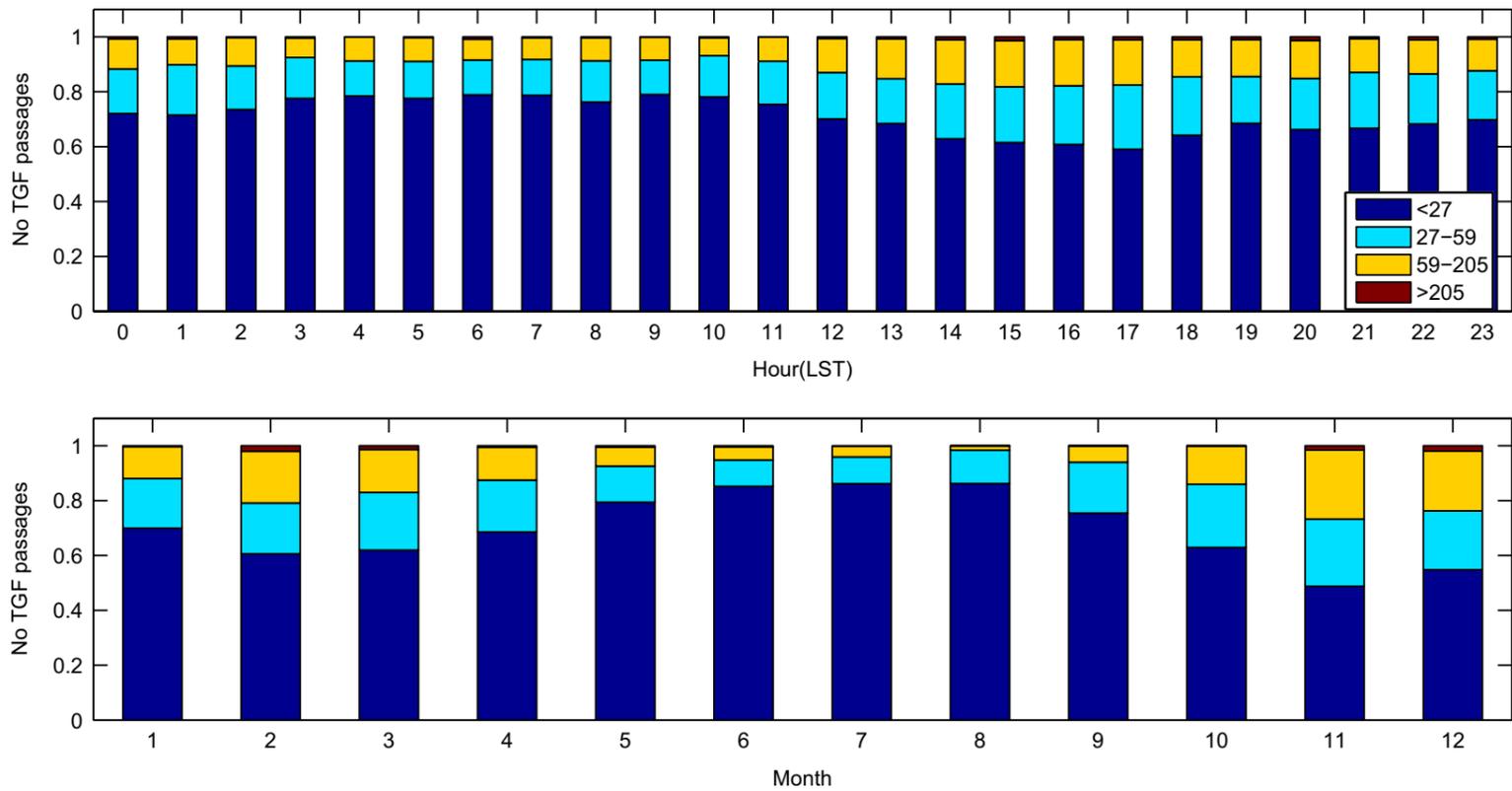


Fig. 7. Number of strokes. Diurnal cycle (top) and monthly distribution (bottom) of the number of strokes of AGILE passages non-detecting TGFs for March 2009–July 2012 period over South America. Values of the four ranges correspond to AGILE passages detecting TGFs (Table 1).

separation in mixed phase, suggesting that strong electrification of mixed phase higher up in the cloud (elevated charge mechanism MacGorman et al., 1989) plays an important role on TGF production. Moreover, more than 50% of TGFs were detected under CAPE values above 3400 J/kg (Table 1, that are not very typical conditions even in the tropics). The fact that TGFs are detected above large number of strokes and storms may be explained by greater chance that one of the thunderstorms is in a stage of development favourable to produce a TGF. Large flash rates also increase the probability to detect a TGF. Finally, the altitude of thunderstorms seems to be important but not promordial for TGF production because 50% of TGFs were detected above CTA >16 km (Table 1) but it is a quite typical situation in the tropics, comprising 60% of the AGILE passages without detected TGFs analysed.

The reason for the small number of TGFs detected outside the tropics would be due to all these conditions that seem to be

preferred for TGF production and that are very rare outside tropical regions, together with analysis presented before showing that the same AGILE satellite over Europe would need ~113 days of exposure to observe a TGF.

4. Conclusions

This paper presents a detailed study of the meteorological conditions conducive to TGF production over South America. The global analysis of RHESSI and AGILE TGFs and WWLLN strokes in 2009–2012 period has revealed:

1. A new weak TGF production region over the tropical Pacific.
2. The TGF/lightning ratio for AGILE TGFs and RHESSI TGFs is different than expected compared to Fuschino et al. (2011) for

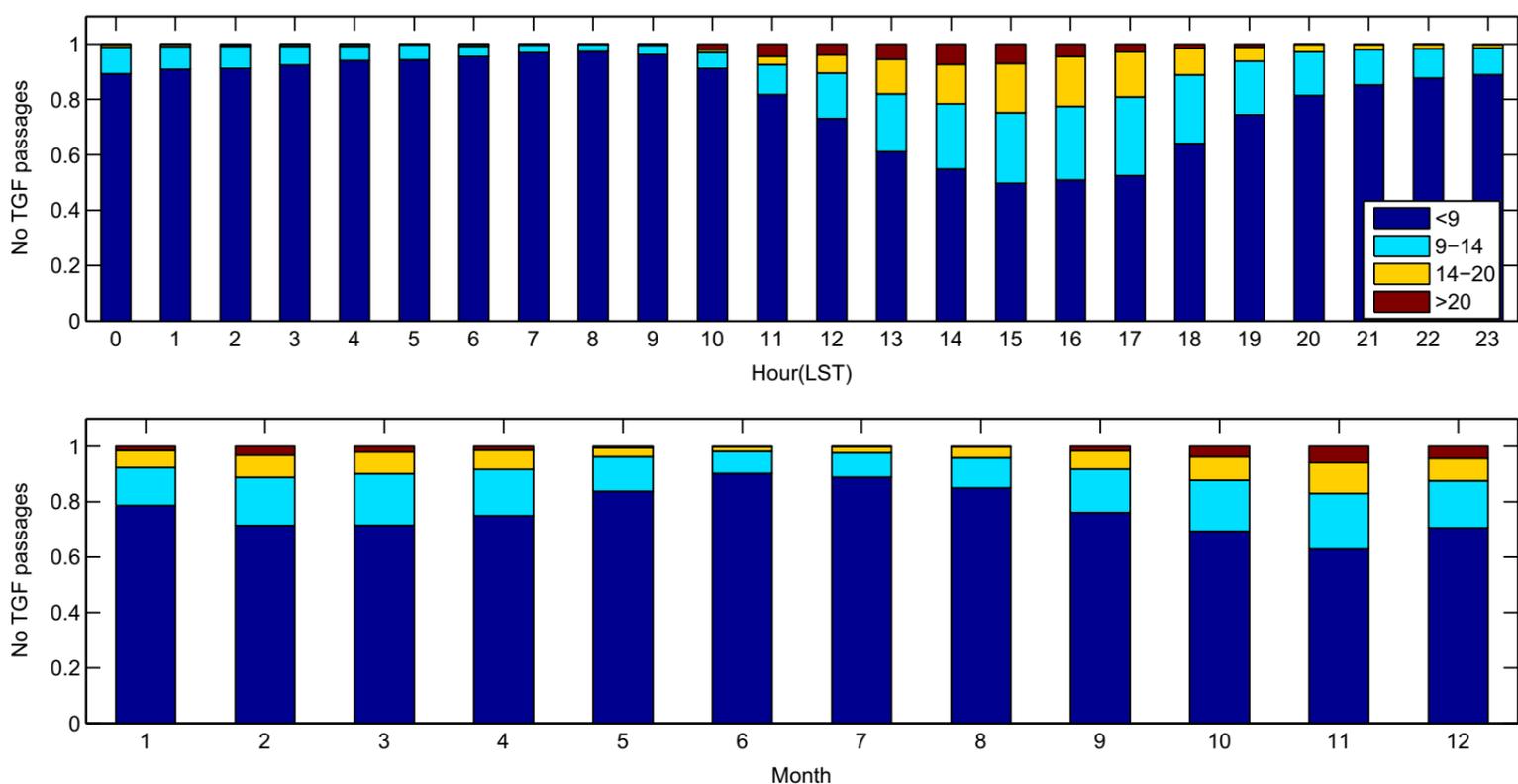


Fig. 8. Number of storms. Diurnal cycle (top) and monthly distribution (bottom) of the number of storms of AGILE passages non-detecting TGFs for March 2009–July 2012 period over South America. Values of the four ranges correspond to AGILE passages detecting TGFs (Table 1).

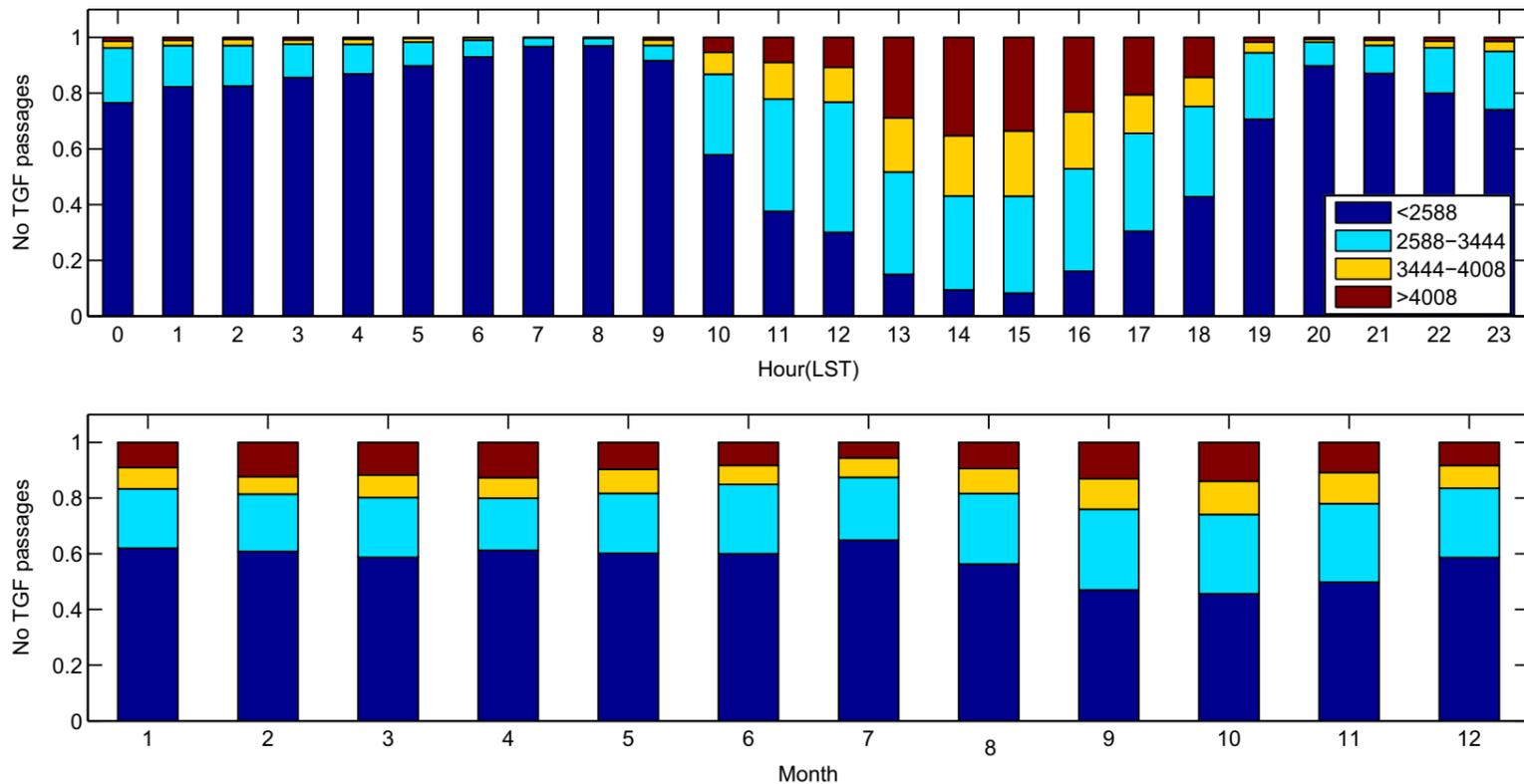


Fig. 9. CAPE. Diurnal cycle (top) and monthly distribution (bottom) of the CAPE of AGILE passages non-detecting TGFs for March 2009–July 2012 period over South America. Values of the four ranges correspond to AGILE passages detecting TGFs (Table 1).

AGILE and Smith et al. (2010) for RHESSI being smaller in America than in Africa and Asia. However these ratios are calculated with a different method and refers mostly to cloud to ground strokes.

3. Geographical occurrence of TGFs in the three main lightning chimneys follows very well the ITCZ seasonal movement as previously reported (Splitt et al., 2010).
4. TGFs exhibit diurnal cycles in all continental regions higher in afternoon suggesting that solar irradiance and consequently CAPE and convective thunderstorms help to get the right conditions for TGF production.

For the analysis of AGILE passages detecting and non-detecting TGFs in the 2009–2012 period over South America we can conclude that:

1. AGILE TGFs are detected under rare conditions of flash rate, number of storms, CAPE and CTC. These conditions occurs more often in the afternoons and months for which the ITCZ is crossing AGILE coverage region.
2. Higher CAPE values suggest that electrification processes related to vigorous updrafts are important for production of lightning associated to TGFs.
3. High number of thunderstorms below the AGILE satellite increases the probability to have a thunderstorm in the most favorable stage of its life cycle (if there is any) for TGF production. Also, a high number of flashes below the AGILE satellite increases the probability to detect a TGF.
4. Cloud Top Altitude (Williams et al., 2006) is not the main parameter for TGF detection at satellite altitude. However it seems important together with high CAPE and lightning activity.

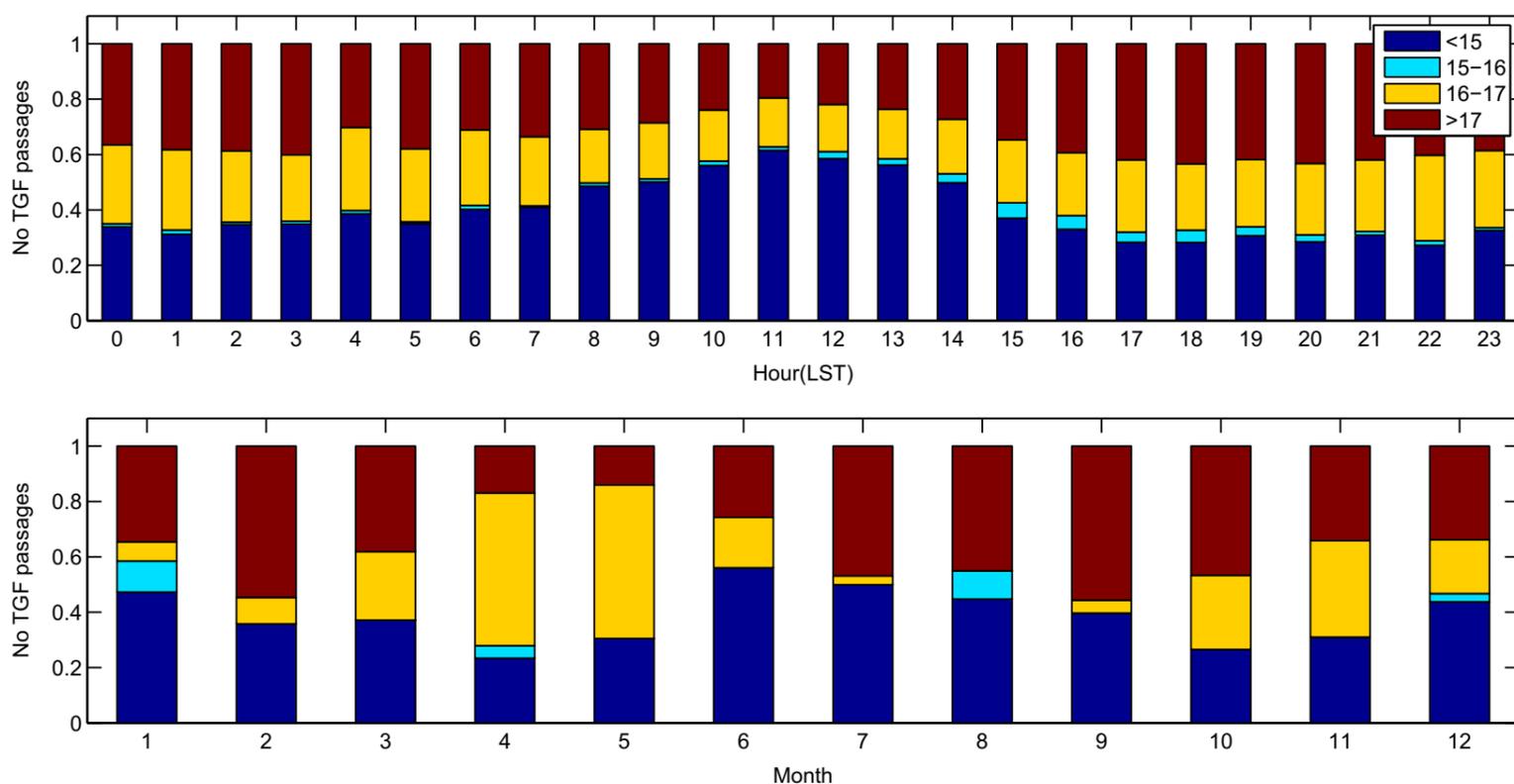


Fig. 10. Cloud Top Altitude. Diurnal cycle (top) and monthly distribution (bottom) of the CTA of AGILE passages non-detecting TGFs for March 2009 - July 2012 period over South America. Values of the four ranges correspond to AGILE passages detecting TGFs (Table 1).

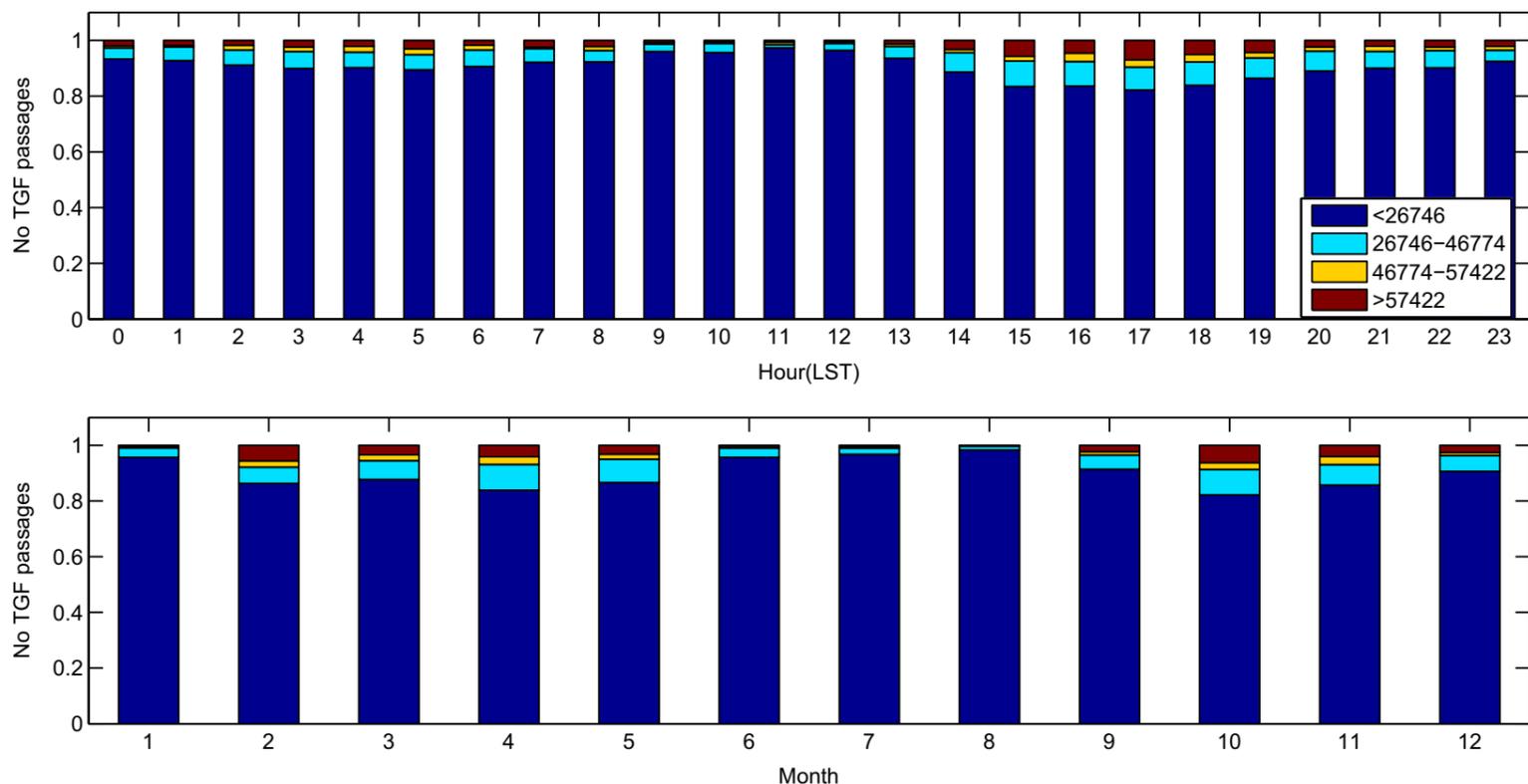


Fig. 11. Cloud Top Coverage. Diurnal cycle (top) and monthly distribution (bottom) of the CTC of AGILE passages non-detecting TGFs for March 2009–July 2012 period over South America. Values of the four ranges correspond to AGILE passages detecting TGFs (Table 1).

5. Preferences for TGF production and comparison of South America and Europe annual stroke densities show that TGF detection in mid-latitudes would be quite difficult with a detector equivalent to MCAL. It would need ~ 113 days of exposure to observe a TGF.

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